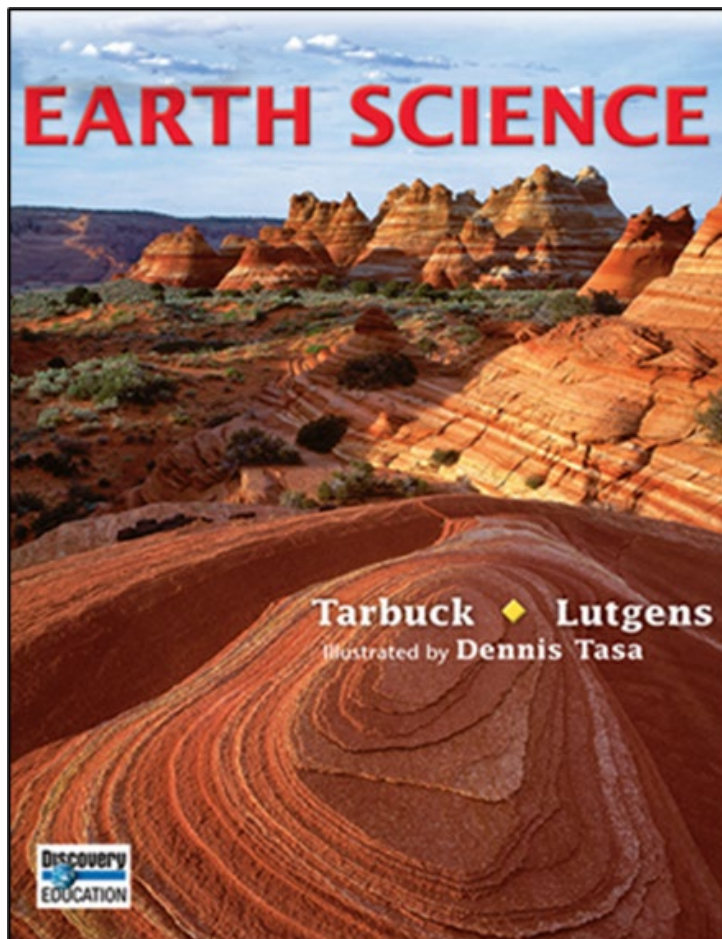


A Correlation of



To the
Oklahoma
2020 Academic Standards for Science
Earth and Space Science

**A Correlation of Earth Science ©2017
To the
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Introduction

This document demonstrates how **Earth Science ©2017** supports the Oklahoma 2020 Academic Standards for Science: Earth and Space Science. Correlation page references are to the Student and Teacher Editions and cited at the page level.

Engage in a journey of observation, explanation, and participation with **Earth Science!**

Renowned authors Edward Tarbuck and Frederick Lutgens invite students on a journey of observation, explanation, and participation in the study of Earth's processes. An accessible writing style combined with digital support create a fresh new program that leads your diverse classroom on a path to discovery. Detailed illustrations by Dennis Tasa provide students with a comprehensive and immersive look at the science behind our planet.

The ©2017 edition of Earth Science features support for the **Next Generation Science Standards** and STEM activities, as well as enhanced resources for both students and teachers:

21st Century Skills: Each chapter of Earth Science an activity geared toward developing one or 21st Century skills. All of these activities task students to capture what they are learning in the science classroom and apply their knowledge to solving real-life problems in order to encourage productive, thoughtful members of the 21st century world.

STEM Activities: STEM activities support the implementation of the engineering process in an engaging and hands-on way. Excite students with real-world engineering design problem and hands-on inquiry. These activities promote higher-order critical thinking skills and result in improved student student performance. Teachers are provided with point-of-use STEM activities and teaching strategies.

Savvas Realize: On savvasrealize.com, you can go digital with online Student Editions and online Teacher Editions, as well as access to editable worksheets.

In addition, Earth Science ©2017 supports the today's diverse classroom with key Spanish resources, including the *Spanish Guided Reading and Study Workbook* and the *Spanish Chapter Tests*.

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Earth's Place in the Universe (ESS1)	
Performance Expectation	
(ES.ESS1.1) Develop a model based on evidence to illustrate the life span of the sun and the role of nuclear fusion in the sun's core to convert matter to energy that eventually reaches Earth in the form of radiation.	SE/TE: Figure 12 Structure of the Sun, 685 The Solar Interior, 689-690 Earth & Space, Solar Variability and Climate Change, 691 Standardized Test Prep (Questions 3-6), 697
Disciplinary Core Ideas	
(ES.ESS1.1.DCI.1) The star called the sun is changing and will burn out over a lifespan of approximately 10 billion years.	SE/TE: How Old Is the Sun?, 690 Figure 10 Lifecycle of a Sun-Like Star, 709
(ES.ESS1.1.DCI.2) Nuclear Fusion processes in the center of the sun release the energy that ultimately reaches Earth as radiation.	SE/TE: The Solar Interior, 689-690
Science and Engineering Practices	
(ES.ESS1.1.SEP.1) Develop and Use Models: Develop and/or use a model (including mathematical and computational) to generate data to support explanations, predict phenomena, analyze systems, and/or solve problems.	SE/TE: Figure 12, Structure of the Sun, 685 Figure 18, Nuclear Fusion, 698 Figure 20, Sunspots, 691,
Crosscutting Concepts	
(ES.ESS1.1.CCC.1) Energy and Matter: Changes of energy and matter in a system can be described in terms of energy and matter flows into, out of, and within that system.	For supporting content, please see: SE/TE: Figure 12, Structure of the Sun, 685 Structure of the Sun, 685-686 The Solar Interior, 689-690
Performance Expectation	
(ES.ESS1.2) Construct an explanation of how the universe formed as a single point and continues to expand based on astronomical evidence of light spectra, motion of distant galaxies, and the composition of matter in the universe.	SE/TE: The Big Bang, 720-721 Figure 23 The Big Bang, 720 Standardized Test Prep, 727
Disciplinary Core Ideas	
(ES.ESS1.2.DCI.1) The study of stars' light spectra and brightness is used to identify compositional elements of stars, their movements, and their distances from Earth.	SE/TE: Characteristics of Stars, 701 Hertzprung-Russell Diagram, 704-706
(ES.ESS1.2.DCI.2) Observations of distant galaxies receding from our own, the measured composition of stars and non-stellar gases, and maps of spectra of the primordial radiation (cosmic microwave background) that still fills the universe are used as evidence to support the explanation of formation.	SE/TE: The Milky Way Galaxy, 715-716 Types of Galaxies, 717-718 The Expanding Universe, 718-719 The Big Bang, 720-721

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(ES.ESS1.2.DCI.3) Other than the hydrogen and helium formed at the time of formation, nuclear fusion within stars produces all atomic nuclei lighter than and including iron, and the process releases electromagnetic energy. Heavier elements are produced when certain massive stars achieve a supernova stage and explode.	SE/TE: Burnout and Death, 710-712 Stellar Remnants, 712-714
(ES.ESS1.2.DCI.4) Atoms of each element emit and absorb characteristic frequencies of light. These characteristics allow identification of the presence of an element, even in microscopic quantities.	SE/TE: Spectroscopy, 676
Science and Engineering Practices	
(ES.ESS1.2.SEP.1) Constructing Explanations: Use evidence (e.g., measurements, observations, patterns) to construct or support an explanation or design a solution to a problem.	SE/TE: Concepts in Action, 696 Reading Checkpoint, 676 24.1 Assessment, 677
Crosscutting Concepts	
(ES.ESS1.2.CCC.1) Matter and Energy: Energy cannot be created or destroyed. It only moves between one place to another, between objects and/or fields, or between systems.	For supporting content, please see: SE/TE: Energy Transfer as Heat, 483-485 What Happens to Solar Radiation, 486-487 The Solar Interior, 689-690
Performance Expectation	
(ES.ESS1.3) Construct an explanation about the process that causes stars to produce elements throughout their life cycle.	For supporting content, please see: SE/TE: Burnout and Death, 710-712 Stellar Remnants, 712-714
Disciplinary Core Ideas	
(ES.ESS1.3.DCI.1) The study of stars' light spectra and brightness is used to identify compositional elements of stars, their movements, and their distances from Earth.	SE/TE: Characteristics of Stars, 701 Measuring Distance to Stars, 702 Stellar Brightness, 703 Hertzprung-Russell Diagram, 704-706 The Expanding Universe, 718-719
(ES.ESS1.3.DCI.2) Other than the hydrogen and helium formed at the time of formation, nuclear fusion within stars produces all atomic nuclei lighter than and including iron, and the process releases electromagnetic energy.	SE/TE: Electromagnetic Radiation, 674-675 Figure 12, Structure of the Sun, 685 Figure 18, Nuclear Fusion, 698 Structure of the Sun, 685-686 The Solar Interior, 689-690 Burnout and Death, 710-712 Stellar Remnants, 712-714

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Science and Engineering Practices	
(ES.ESS1.3.SEP.1) Constructing Explanations: Use evidence (e.g., measurements, observations, patterns) to construct or support an explanation	SE/TE: Earth & Space, Solar Variability and Climate Change, 691 Inquiry: Exploration Lab?, Teacher Demo: Tracking Sun Spots, 692-693
Crosscutting Concepts	
(ES.ESS1.3.CCC.1) Cause and Effect: Empirical evidence is required to differentiate between cause and correlation and make claims about specific causes and effects.	For supporting content, please see: SE/TE: Inquiry: Exploration Lab?, Teacher Demo: Tracking Sun Spots, 692-693 Earth & Space, Solar Variability and Climate Change, 691 Performance-Based Assessment, 696
Performance Expectation	
(ES.ESS1.4) Use mathematical or computational representations to determine patterns that can be used to predict the motion of orbiting objects in the solar system.	SE/TE: Earth-Sun Relationships, 481-482 Ancient Greeks, 615-616 Figure 4, Retrograde Motion, 617 The Birth of Modern Astronomy, 617-621 Table 1, Period of Revolution and Solar Distance of Planets, 618 22.1 Assessment, Math Practice, 621 Motions of Earth, 622-627 Figure 13: Sidereal Day, 623 Inquiry: Exploration Lab: Modeling Synodic and Sidereal Months, 636-637
Disciplinary Core Ideas	
(ES.ESS1.4.DCI.1) The solar system consists of the sun and a collection of objects of varying sizes and conditions-including planets and their moons-that are held in orbit around the sun by its gravitational pull on them.	SE/TE: Earth-Sun Relationships, 481-482 Ancient Greeks, 615-616 Figure 4, Retrograde Motion, 617 The Birth of Modern Astronomy, 617-621 Table 1, Period of Revolution and Solar Distance of Planets, 618 The Solar System, 644 The Planets: An Overview, 645-646 Mercury: The Innermost Planet, 649-650 Venus: The Veiled Planet, 650-651 Mar: The Red Planet, 652-653 Jupiter: Giant Among Planets, 654-655 Saturn: The Elegant Planet, 656-657 Uranus: The Sideways Planet, 658 Neptune: The Windy Planet, 658 Pluto: Dwarf Planet, 659

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(ES.ESS1.4.DCI.2) Kepler's laws describe common features of the motions of orbiting objects, including their elliptical paths around the sun. Orbits may change due to the gravitational effects from, or collisions with, other objects in the solar system.	SE/TE: The Birth of Modern Astronomy, 617-621 Figure 6: Planet Revolution, 618 Table 1, Period of Revolution and Solar Distance of Planets, 618
Science and Engineering Practices	
(ES.ESS1.4.SEP.1) Using mathematics and computational thinking Use mathematical representations of phenomena or design solutions to support and revise explanations.	SE/TE: Earth-Sun Relationships, 481-482 Ancient Greeks, 615-616 Figure 4, Retrograde Motion, 617 The Birth of Modern Astronomy, 617-621 Table 1, Period of Revolution and Solar Distance of Planets, 618 22.1 Assessment, Math Practice, 621 Motions of Earth, 622-627 Figure 13: Sidereal Day, 623 Inquiry: Exploration Lab: Modeling Synodic and Sidereal Months, 636-637
Crosscutting Concepts	
(ES.ESS1.4.CCC.1) Patterns: Mathematical representations are needed to identify some patterns.	For supporting content, please see: SE/TE: The Birth of Modern Astronomy, 617-621 Figure 6: Visual Summary: The Solar System Model Evolves, 619 Motions of the Earth, 622-625 Motions of the Earth-Sun-Moon System, 626-627 Eclipses, 628-629
Performance Expectation	
(ES.ESS1.5) Evaluate evidence in the patterns of the past and current movements of continental and oceanic crust and the theory of plate tectonics to explain the ages of crustal rocks.	SE/TE: Inquiry: Try It!: How Do The Continents Fit Together, 247 Evidence for Continental Drift, 249-251 Rejection of Wegener's Hypothesis, 253 Evidence of Sea-Floor Spreading, 257-260 Earth's Moving Plates, 261-263 Divergent Boundaries, 264 Convergent Boundaries, 265-267 Transform Fault Boundaries, 268 What Causes Plate Motions?, 270 Plate Motion Mechanisms, 271
Disciplinary Core Ideas	
(ES.ESS1.5.DCI.1) Plate tectonics is the unifying theory that explains the past and current movements of the rocks at Earth's surface and provides a framework for understanding its geologic history.	SE/TE: Earth's Moving Plates, 261-263 Divergent Boundaries, 264 Convergent Boundaries, 265-267 Transform Fault Boundaries, 268

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Science and Engineering Practices	
(ES.ESS1.5.SEP.1) Engaging in Argument from Evidence: Evaluate the claims, evidence, and/or reasoning behind currently accepted explanations or solutions to determine the merit of arguments.	For supporting content, please see: SE/TE: 9.2 Assessment, 260 Inquiry: Exploration Lab: Paleomagnetism and the Ocean Floor, 272-273 9 Assessment, Think Critically, 276
Crosscutting Concepts	
(ES.ESS1.5.CCC.1) Stability and Change: Much of science deals with constructing explanations of how things change and how they remain stable.	For supporting content, please see: SE/TE: Earth's Moving Plates, 261-263 Divergent Boundaries, 264 Convergent Boundaries, 265-267 Transform Fault Boundaries, 268 Stability, 514-515
Performance Expectation	
(ES.ESS1.6) Apply scientific reasoning and evidence from ancient Earth materials, meteorites, and other planetary surfaces to construct an account of changes in Earth's formation and early history.	SE/TE: Uniformitarianism, 336 Relative Dating, 337-340 Correlation, 340-341 Types of Fossils, 342-342 The Fossil Record, 344-346 Radiometric Dating, 348-349 Dating with Carbon-14, 350 Radiometric Dating of Sedimentary Rock, 350-351 Lunar History, 633-634 The Planets: An Overview, 645-646 Formation of the Solar System, 647-648
Disciplinary Core Ideas	
(ES.ESS1.6.DCI.1) Although active geologic processes, such as plate tectonics and erosion, have destroyed or altered most of the very early rock record on Earth, other objects in the solar system, such as lunar rocks, asteroids, and meteorites, have changed little over billions of years. Studying these objects can provide information about Earth's formation and early history.	SE/TE: The Lunar Surface, 631-632 Lunar History, 633-634 Formation of the Solar System, 647-648 Asteroids, 660 Comets, 661-662 Meteoroids, 663-664 Earth & Space: Is Earth on a Collision Course?, 665
(ES.ESS1.6.DCI.2) Spontaneous radioactive decays follow a characteristic exponential decay law. Nuclear lifetimes allow radiometric dating to be used to determine the ages of rocks and other materials.	SE/TE: What is Radioactivity?, 347-348 Radiometric Dating, 348-349 Dating with Carbon-14, 350 Radiometric Dating of Sedimentary Rock, 350-351

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Science and Engineering Practices	
(ES.ESS1.6.SEP.1) Constructing Explanations: Apply scientific reasoning, theory, and/or models to link evidence to the claims to assess the extent to which the reasoning and data support the explanation or conclusion.	SE/TE: 12.2 Assessment, 346 Inquiry: Apply It?, 349 12.3 Assessment, 351 Earth and History: Dating with Tree Rings, 352 Chapter 12 Assessment, 360
Crosscutting Concepts	
(ES.ESS1.6.CCC.1) Stability and Change: Much of science deals with constructing explanations of how things change and how they remain stable.	For supporting content, please see: SE/TE: Earth's Moving Plates, 261-263 Divergent Boundaries, 264 Convergent Boundaries, 265-267 Transform Fault Boundaries, 268 Stability, 514-515
Earth Systems (ESS2)	
Performance Expectation	
(ES.ESS2.1) Develop a model to illustrate how Earth's internal and surface processes operate at different scales of space and time to form continental and ocean-floor features.	SE/TE: Exploring the Ocean Floor, 254-255 The Processes of Sea-Floor Spreading, 256-257 Evidence of Sea-Floor Spreading, 257-260 Earth's Moving Plates, 261-263 Divergent Boundaries, 264 Convergent Boundaries, 265-267 Transform Fault Boundaries, 268 Earth & Its Systems: Plate Tectonics into the Future, 269 What Causes Plate Motions?, 270 Plate Motion Mechanisms, 271 Inquiry: Exploration Lab? Paleomagnetism and the Ocean Floor, 272-273 How Earth Works: Effects of Volcanoes, 298-299 Folds, 312-313 Faults, 314-315 Types of Mountains, 316-317 Plateaus, Domes, and Basins, 318-319 Inquiry: Quick Lab?: Rates of Mountain Building, 323 Earth & Its Systems: Mountain Building Away from Plate Margins, 326-327
Disciplinary Core Ideas	
(ES.ESS2.1.DCI.1) Earth's systems, being dynamic and interacting, cause feedback effects that can increase or decrease the original changes.	SE/TE: What is a System, 18 Earth as a System, 19 Earth & Its Systems: Plate Tectonics into the Future, 269 Precambrian Life, 367-368 Natural Processes that Change Climate, 600-601 Human Impact on Climate, 602-603

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(ES.ESS2.1.DCI.2) Plate tectonics is the unifying theory that explains the past and current movements of rocks at Earth's surface and provides a framework for understanding its geologic history.	SE/TE: Earth's Moving Plates, 261-263 Divergent Boundaries, 264 Convergent Boundaries, 265-267 Transform Fault Boundaries, 268 Earth & Its Systems: Plate Tectonics into the Future, 269 What Causes Plate Motions?, 270 Plate Motion Mechanisms, 271 Relative Dating, 337-340 Correlation, 340-341 Types of Fossils, 342-342 The Fossil Record, 344-346 Radiometric Dating, 348-349 Dating with Carbon-14, 350 Radiometric Dating of Sedimentary Rock, 350-351
(ES.ESS2.1.DCI.3) Plate movements are responsible for most continental and ocean-floor features and for the distribution of most rocks and minerals within the Earth's crust.	SE/TE: Exploring the Ocean Floor, 254-255 The Processes of Sea-Floor Spreading, 256-257 Evidence of Sea-Floor Spreading, 257-260 Earth's Moving Plates, 261-263 Divergent Boundaries, 264 Convergent Boundaries, 265-267 Transform Fault Boundaries, 268 Earth & Its Systems: Plate Tectonics into the Future, 269 Inquiry: Exploration Lab? Paleomagnetism and the Ocean Floor, 272-273
Science and Engineering Practices	
(ES.ESS2.1.SEP.1) Developing and Using Models: Develop a model based on evidence to illustrate the relationships between systems or components of a system.	SE/TE: Earth & Its Systems: Plate Tectonics into the Future, 269 Inquiry: Exploration Lab? Paleomagnetism and the Ocean Floor, 272-273 Standardized Test Prep, 277 TE only: Building Science Skills: Use Models, 257 A Teacher Demo: A Convergent Model, 262
Crosscutting Concepts	
(ES.ESS2.1.CCC.1) Scale Proportion and Quantity: Some systems can only be studied indirectly as they are too small, too large, too fast, or too slow to observe directly.	For supporting content, please see: SE/TE: Structure of the Time Scale, 354-355 Precambrian Earth, 364-366 Cambrian Period, 369-370 Inquiry: Exploration Lab?: Modeling the Geologic Time Scale, 386-387

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Performance Expectation	
(ES.ESS2.2) Analyze geoscience data to make the claim that one change to Earth's surface can create feedbacks and interactions that cause changes to other Earth's systems.	SE/TE: What is a System, 18 Earth as a System, 19-20 People and the Environment, 20-21 Folds, 312-313 Faults, 314-315 Types of Mountains, 316-317 Plateaus, Domes and Basins, 318-319 Convergent Boundary Mountains, 320-322 Divergent Boundary Mountains, 323 Non-Boundary Mountains, 323 Continental Accretion, 324-325 Inquiry-Quick Lab: Rates of Mountain Building, 323 Earth & Its Systems: Mountain Building Away from Plate Margins, 326-327 How Earth Works: Effects of Volcanoes, 298-299
Disciplinary Core Ideas	
(ES.ESS2.2.DCI.1) Earth's systems, being dynamic and interacting, cause feedback effects that can increase or decrease the original changes.	SE/TE: Erosion, 164 Sediment Transport, 164-165 Deposition, 166 Stream Valleys, 167-168 Wind Erosion, 203-204 Wind Deposits, 204-205 How Earth Works: Erosion, 208-209 Deformation of Rock, 308-309 Types of Stress, 309 Isostasy, 310-311
(ES.ESS2.2.DCI.2) The foundation for Earth's global climate system is the electromagnetic radiation from the sun, as well as its reflection, absorption, storage, and redistribution among the atmosphere, ocean, and land systems, and this energy's re-radiation into space.	SE/TE: Energy Transfer as Heat, 483 What Happens to Solar Radiation?, 486-487 Factors That Determine Climate, 588-590 Figure 2, Earth's Major Climate Zones, 589 Inquiry: Quick Lab: Observing How Land and Water Absorb and Release Energy, 590
Science and Engineering Practices	
(ES.ESS2.2.SEP.1) Analyzing and Interpreting Data: Analyze data using tools, technologies, and/or models in order to make valid and reliable scientific claims.	SE/TE: Figure 3, Primary Pollutants in the Atmosphere, 478 Figure 4, Atmospheric Pressure vs. Altitude, 479 Figure 15, Mean Monthly Temperatures for Vancouver and Winnipeg, 489 Table 1: Variation in Annual Mean Temperature Range with Latitude, 489 Inquiry: Exploration Lab: Heating Land and Water, 496-497 Chapter 17 Assessment: Think Critically, 500 Standardized Test Prep, 501

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Crosscutting Concepts	
(ES.ESS2.2.CCC.1) Stability and Change: Feedback (negative or positive) can stabilize or destabilize a system.	For supporting content, please see: SE/TE: Earth's Moving Plates, 261-263 Divergent Boundaries, 264 Convergent Boundaries, 265-267 Transform Fault Boundaries, 268 Surface Circulation, 448-451 Deep-Ocean Circulation, 451-453 Stability, 514-515 Natural Processes that Change Climate, 600-601 Human Impact on Climate, 602-603
Performance Expectation	
(ES.ESS2.3) Develop a model based on evidence of Earth's interior to describe the cycling of matter by thermal convection.	SE/TE: The Process of Sea-Floor Spreading, 256-257 Figure 10: Sea-Floor Spreading and Subduction, 257 What Causes Plate Motions?, 270-271 Earth & Its Systems: The Carbon Cycle, 85
Disciplinary Core Ideas	
(ES.ESS2.3.DCI.1) Evidence from deep probes and seismic waves, reconstructions of historical changes in Earth's surface features, its magnetic field, and an understanding of physical and chemical processes lead to a model of Earth with a hot but solid inner core, a liquid outer core, a solid mantle and crust.	SE/TE: Seismic Waves, 222-224 Recording Seismic Waves, 224 Measuring Earthquakes, 225-226 Locating Earthquakes, 226-227 Layers Defined by Composition, 234-235 Figure 15: Paths of Seismic Waves, 233 Discovering Earth's Layers, 236 Figure 18: Earth's Interior Showing P and S Wave Paths, 236 Discovering Earth's Composition, 237
(ES.ESS2.3.DCI.2) Motions of the mantle and its plates occur primarily through thermal convection, which involves the cycling of matter due to the outward flow of energy from Earth's interior and gravitational movement of denser materials toward the interior.	SE/TE: Layers Defined by Composition, 234-235 The Process of Sea-Floor Spreading, 256-257 Figure 10: Sea-Floor Spreading and Subduction, 257 Convergent Boundaries, 265-267 What Causes Plate Motions?, 270-271 Plate Motion Mechanisms, 271
(ES.ESS2.3.DCI.3) The radioactive decay of unstable isotopes continually generates new energy within Earth's crust and mantle, providing the primary source of the heat that drives mantle convection.	SE/TE: What causes Plate Motions?, 270 What Is Radioactivity?, 347-348
(ES.ESS2.3.DCI.4) Plate tectonics can be viewed as the surface expression of mantle convection.	SE/TE: Plate Motion Mechanisms, 271

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Science and Engineering Practices	
(ES.ESS2.3.SEP.1) Developing and Using Models: Develop a model based on evidence to illustrate the relationships between systems or components of a system.	<p>SE/TE: Figure 3: Matching Mountain Ranges, 250 Figure 4: Glacier Evidence, 251 Visual Summary: Breakup of Pangaea, 252</p> <p>TE only: Build Science Skills: Use Models, 250 Teacher Demo: A Convergent Model</p>
Crosscutting Concepts	
(ES.ESS2.3.CCC.1) Energy and Matter: Energy drives the cycling of matter within and between systems.	<p>For supporting content, please see: SE/TE: What Causes Plate Motions?, 270 What Is Radioactivity?, 347-348 Figure 23: Whole Mantle Convection, 271</p>
Performance Expectation	
(ES.ESS2.4) Analyze and interpret data to explore how variations in the flow of energy into and out of Earth's systems causes changes to the atmosphere and climate.	<p>SE/TE: Energy Transfer as Heat, 483-485 What Happens to Solar Radiation, 486-487 How Earth Works: Earth's Atmosphere, 494-495 Inquiry: Exploration Lab: Heating Land and Water, 496-497 Water's Changes of State, 504-506 Factors Affecting Wind, 534-536 El Nino and La Nina, 546-547 Powered By The Sun, 588 Inquiry Lab: Quick Lab: Observing How Land and Water Absorb and Release Energy, 590 Natural Changes That Change Climate, 600-601 Human Impact on Climate, 602-603</p>
Disciplinary Core Ideas	
(ES.ESS2.4.DCI.1) The geological record shows that changes to global and regional climate can be caused by interactions among changes in the sun's energy output or Earth's orbit, tectonic events, ocean circulation, volcanic activity, glaciers, vegetation, and human activities. These changes can occur on a variety of time scales from sudden (e.g., volcanic ash clouds) to intermediate (ice ages) to very long-term (tectonic cycles).	<p>SE/TE: Types of Glaciers, 188-189 Ancient Climates, 250-251 Natural Processes That Change Climate, 600-601</p> <p>TE only: Earth Science Refresher, 586C-586D</p>

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(ES.ESS2.4.DCI.2) The foundation for Earth’s global climate system is the electromagnetic radiation from the sun, as well as its reflection, absorption, storage, and redistribution among the atmosphere, ocean, and land systems, and this energy’s re-radiation into space.	SE/TE: Energy Transfer as Heat, 483-485 What Happens to Solar Radiation, 486-487 How Earth Works: Earth’s Atmosphere, 494-495 Water’s Changes of State, 504-506 Factors Affecting Wind, 534-536 El Nino and La Nina, 546-547 Powered by the Sun, 588 Natural Changes That Change Climate, 600-601 Human Impact on Climate, 602-603
(ES.ESS2.4.DCI.3) Cyclical changes in the shape of Earth’s orbit around the sun, together with changes in the tilt of the planet’s axis of rotation, both occurring over hundreds of thousands of years, have altered the intensity and distribution of sunlight falling on the Earth. These phenomena cause a cycle of ice ages and other changes in climate.	SE/TE: Natural Changes That Change Climate, 600-601 Quaternary Period, 384-385
Science and Engineering Practices	
(ES.ESS2.4.SEP.1) Analyzing and Interpreting Data: Analyze data using computational models in order to make valid and reliable scientific claims.	SE/TE: Table 1: Water Vapor Needed for Saturation, 506 Figure 3, Relative Humidity, 507 Inquiry: Exploration Lab: Measuring Humidity, 524-525 Chapter 18 Assessment: Math Skills, 528 Inquiry: Exploration Lab: Observing Wind Patterns, 550-551 Chapter 19 Assessment, Math Skills, 554
Crosscutting Concepts	
(ES.ESS2.4.CCC.1) Cause and Effect: Empirical evidence is required to differentiate between cause and correlation and make claims about specific causes and effects.	For supporting content, please see: SE/TE: Natural Processes That Change Climate, 600-601 Figure 15: Change in CO2 Levels, 602 Inquiry: Exploration Lab: Human Impact on Climate and Weather, 606-607 Chapter 21 Assessment, 610

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Performance Expectation	
(ES.ESS2.5) Plan and conduct investigations of how the structure and resulting properties of water interact with the Earth’s materials and surface processes.	For supporting content, please see: SE/TE: Mechanical Weathering, 126-128 Water Cycle, 158-159 Streamflow, 160 Sediment Transport, 164-165 Deposition, 166-167 Stream Valleys, 167-168 Floods and Flood Control, 168-169 Drainage Basins, 170 Movement and Distribution of Water Underground, 171-172 Springs, 172-173 Types of Glaciers, 188-189 How Glaciers Move, 190-191 Glacial Erosion, 192 Landforms Formed by Glacial Erosion, 193-194 Glacial Deposition, 194 Moraines, Outwash Plains, and Kettles, 195-197 Glaciers of the Ice Age, 197-198 Water in Deserts, 200-202 Origin of Magma, 280-281 Why Temperatures Vary, 488-492 Water’s Changes of State, 504-506 Humidity, 506-509 Fog, 520 How Precipitation Forms, 520-522 Forms of Precipitation, 522 Natural Processes that Change Climate, 600-601
Disciplinary Core Ideas	
(ES.ESS2.5.DCI.1) The abundance of liquid water on Earth’s surface and its unique combination of physical and chemical properties are central to the planet’s dynamics. These properties include water’s exceptional capacity to absorb, store, and release large amounts of energy; transmit sunlight; expand upon freezing; dissolve and transport materials; and lower the viscosities and melting points of rocks.	SE/TE: Mechanical Weathering, 126-128 Water Cycle, 158-159 Streamflow, 160 Sediment Transport, 164-165 Deposition, 166-167 Stream Valleys, 167-168 Floods and Flood Control, 168-169 Origin of Magma, 280-281 Why Temperatures Vary, 488-492 Water’s Changes of State, 504-506 Humidity, 506-509 Fog, 520 How Precipitation Forms, 520-522 Forms of Precipitation, 522 Natural Processes that Change Climate, 600-601

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(ES.ESS2.5.SEP.1) Planning and Conducting Investigations: Plan and conduct an investigation individually and collaboratively to produce data to serve as the basis for evidence.	SE/TE: Inquiry: Try It!: What Causes Weathering, 125 Inquiry: Exploration Lab: Effect of Temperature on Chemical Weathering, 150-151 Inquiry: Try It!: How do Local Bodies of Water Affect Your Community, 157 Inquiry: Exploration Lab: Investigating the Permeability of Soils, 181 Inquiry: Exploration Lab: Measuring Humidity, 524-525
Crosscutting Concepts	
(ES.ESS2.5.CCC.1) Structure and Function: The functions and properties of natural and designed objects and systems can be inferred from their overall structure, the way their components are shaped and used, and the molecular substructures of their various materials.	For supporting content, please see: SE/TE: Mechanical Weathering, 126-128 Types of Chemical Bonds, 40-43 How Minerals Form, 45-46 Mineral Groups, 47-49 Distinctive Properties of Minerals & Table 1, 54-55
Performance Expectation	
(ES.ESS2.6) Develop a quantitative model to describe the cycling of carbon among the hydrosphere, atmosphere, geosphere, and biosphere.	For supporting content, please see: SE/TE: Isotopes, 38 Mineral Groups, 47-49 Earth & Its Systems: The Carbon Cycle, 85 Earth's Blanket of Air, 110
Disciplinary Core Ideas	
(ES.ESS2.6.DCI.1) Gradual atmospheric changes were due to plants and other organisms that captured carbon dioxide and released oxygen.	SE/TE: The Cambrian Earth, 364-366 Precambrian Life, 367-368 How Earth Works: Earth's Atmosphere, 495-496
(ES.ESS2.6.DCI.2) Changes in the atmosphere due to human activity have increased carbon dioxide concentrations and thus affect climate.	SE/TE: Factors that Determine Climate, 588-591 Human Impact on Climate, 602-603 Figure 15: Change in CO ₂ Levels, 602 Inquiry: Exploration Lab: Human Impact on Climate and Weather, 606-607
Science and Engineering Practices	
(ES.ESS2.6.SEP.1) Developing and Using Models: Develop a model based on evidence to illustrate the relationships between systems or components of a system.	SE/TE: TE: Modeling Exfoliation, 128 How Earth Works: Earth's Atmosphere, 495-496 TE only: Modeling Exfoliation, 128 Modeling Humid Climates, 596

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Crosscutting Concepts	
(ES.ESS2.6.CCC.1) Energy and Matter: Energy drives the cycling of matter within and between systems.	For supporting content, please see: SE/TE: Earth & Its Systems: The Carbon Cycle, 85 The Water Cycle, 158-159 Figure 2: The Water Cycle, 159 How Earth Works: Earth's Atmosphere, 495-496
Performance Expectation	
(ES.ESS2.7) Engage in argument from evidence for how the simultaneous co-evolution of Earth's systems and life on Earth led to periods of stability and change over geologic time.	For supporting content, please see: SE/TE: The Cambrian Earth, 364-366 Precambrian Life, 367-368 Cambrian Period, 369-370 Ordovician Period, 371 Silurian Period, 372 Devonian Period, 372-373 Carboniferous Period, 374 Permian Period, 375-376 13.2 Assessment: Big Idea, 376 Triassic Period, 377-378 Jurassic Period, 379-380 Cretaceous Period, 380 Cretaceous Extinction, 381 The Rise of Mammals, 382-383 Tertiary Period, 383 Quaternary Period, 384-385 Chapter 13 Assessment, 390
Disciplinary Core Ideas	
(ES.ESS2.7. DCI.1) Gradual atmospheric changes were due to plants and other organisms that captured carbon dioxide and released oxygen.	SE/TE: The Cambrian Earth, 364-366 Precambrian Life, 367-368 Cambrian Period, 369-370
(ES.ESS2.7. DCI.2) The many dynamic and delicate feedback mechanisms between the biosphere and other Earth systems cause a continual co-evolution of Earth's surface and the life that exists on it.	SE/TE: Earth's Major Spheres, 7-9 What is a System, 18 The Water Cycle, 158-159 The Cambrian Earth, 364-366 Precambrian Life, 367-368 Cambrian Period, 369-370
Science and Engineering Practices	
(ES.ESS2.7. SEP.1) Engaging in Argument from Evidence: Construct an oral and written argument or counter-argument based on data and evidence.	For supporting content, please see: SE/TE: 13.2 Assessment, 376 Chapter 13 Assessment, 390

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Crosscutting Concepts	
(ES.ESS2.7.CCC.1) Stability and Change: Much of science deals with constructing explanations of how things change and how they remain stable.	For supporting content, please see: SE/TE: The Cambrian Earth, 364-366 Precambrian Life, 367-368 Cambrian Period, 369-370 Ordovician Period, 371 Silurian Period, 372 Devonian Period, 372-373 Carboniferous Period, 374 Permian Period, 375-376 Triassic Period, 377-378 Jurassic Period, 379-380 Cretaceous Period, 380 Cretaceous Extinction, 381 The Rise of Mammals, 382-383 Tertiary Period, 383 Quaternary Period, 384-385 Stability, 514-515
Earth and Human Activities (ESS3)	
Performance Expectation	
(ES.ESS3.1) Construct an explanation based on evidence for how the availability of natural resources, occurrence of natural hazards, and changes in climate effect human activity.	SE/TE: Renewable and Nonrenewable Resources, 94-95 Fossil Fuels, 95-96 Tar Sands and Oil Shale, 97 Solar Energy, 102 Nuclear Energy, 103-104 Wind Energy, 104 Hydroelectric Power, 105 Geothermal Energy, 105-106 Tidal Power, 106-107 The Water Planet, 108-109 Earth's Blanket of Air, 110 Land Resources, 111-112 Keeping Water Clean and Safe, 114 Protecting the Air, 114-115 Caring for Land Resources, 115-116 Floods and Flood Control, 168-169 Causes of Earthquake Damage , 228-230 Reducing Earthquake Damage, 231-232 8.3 Assessment, 232 How Earth Works: Winds and Storms, 578-579 Chapter 20 Assessment, 584 Human Impact on Climate, 602-603 Inquiry: Exploration Lab: Human Impact on Climate and Weather, 606-607

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Disciplinary Core Ideas	
(ES.ESS3.1.DCI.1) Resource availability has guided the development of human society.	For supporting content, please see: SE/TE: The Water Planet, 108-109 Earth's Blanket of Air, 110 Land Resources, 111-112 Keeping Water Clean and Safe, 114 Protecting the Air, 114-115 Caring for Land Resources, 115-116 Chapter 4 Assessment, 122
(ES.ESS3.1.DCI.2) Natural hazards and other geologic events have shaped the course of human history; they have significantly altered the sizes of human populations and have driven human migrations.	For supporting content, please see: SE/TE: People and the Environment, 20-21 Environmental Problems, 21-22 Floods and Flood Control, 168-169 Causes of Earthquake Damage, 228-230 Reducing Earthquake Damage, 231-232 8.3 Assessment, 232 Volcanic Hazards, 294 How Earth Works: Effects of Volcanoes, 298-299 Tornadoes, 573-574 Hurricanes, 575-577 How Earth Works: Winds and Storms, 578-579 Chapter 20 Assessment, 584
Science and Engineering Practices	
(ES.ESS3.1.SEP.1) Constructing Explanations: Construct an explanation based on valid and reliable evidence obtained from a variety of sources (including students' own investigations, models, theories, simulations, peer review) and the assumption that theories and laws that describe the natural world operate today as they did in the past and will continue to do so in the future.	SE/TE: The Process of Science, 23 Scientific Knowledge, 24 10.1 Assessment, 285 10.2 Assessment, 294 Chapter 10 Assessment, 304 Uniformitarianism, 336 Inquiry: Exploration Lab: Middle-Latitude Cyclones, 580-581 Inquiry: Exploration Lab: Human Impact on Climate and Weather, 606-607 Chapter 21 Assessment, 610
Crosscutting Concepts	
(ES.ESS3.1.CCC.1) Cause and Effect: Empirical evidence is required to differentiate between cause and correlation and make claims about specific causes and effects.	For supporting content, please see: SE/TE: 21.1 Assessment, 591 Inquiry: Exploration Lab: Human Impact on Climate and Weather, 606-607

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Performance Expectation	
(ES.ESS3.2) Evaluate competing design solutions for developing, managing, and utilizing energy and mineral resources based on cost-benefit ratios on large and small scales.	For supporting content, please see: SE/TE: Try It!: How Can You Determine the Resources You Use?, 93 Earth & Its Resources: Bingham Canyon, Utah: The Largest Open-Pit Mine, 117 Inquiry: Exploration Lab: Finding the Product that Best Conserves Resources, 118-119 Chapter 4 Assessment, 122 Earth and Human Activity: Science and Engineering Practices: Designing Solutions: Design to Reduce Waste, 729
Disciplinary Core Ideas	
(ES.ESS3.2.DCI.1) All forms of energy production and other resource extraction have associated economic, social, environmental, and geopolitical costs and risks as well as benefits. New technologies and social regulations can change the balance of these factors.	SE/TE: Try It!: How Can You Determine the Resources You Use?, 93 Renewable and Nonrenewable Resources, 94-95 Fossil Fuels, 95-96 Tar Sands and Oil Shale, 97 Formation of Mineral Deposits, 98-100 Nonmetallic Mineral Resources, 100-101 4.1 Assessment, 101 Solar Energy, 102 Nuclear Energy, 103-104 Wind Energy, 104 Hydroelectric Power, 105 Geothermal Energy, 105-106 Inquiry: Apply It, 106 Tidal Power, 106-107 The Water Planet, 108-109 Earth's Blanket of Air, 110 Land Resources, 111-112 Keeping Water Clean and Safe, 114 Protecting the Air, 114-115 Caring for Land Resources, 115-116 Inquiry: Exploration Lab: Finding the Product that Best Conserves Resources, 118-119

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(ES.ESS3.2.DCI.2) When evaluating solutions, it is important to take into account a range of constraints, including cost, safety, reliability, and aesthetics, and to consider social, cultural, and environmental impacts.	SE/TE: Try It!: How Can You Determine the Resources You Use?, 93 4.1 Assessment, 101 Wind Energy, 104 Hydroelectric Power, 105 Geothermal Energy, 105-106 Inquiry: Apply It, 106 Tidal Power, 106-107 Caring for Land Resources, 115-116 Chapter 4 Assessment, 122
Science and Engineering Practices	
(ES.ESS3.2.SEP.1) Engaging in Argument from Evidence: Evaluate competing design solutions to a real-world problem based on scientific ideas and principles, empirical evidence, and logical arguments regarding relevant factors (e.g. economic, societal, environmental, ethical considerations).	For supporting content, please see: SE/TE: Try It!: How Can You Determine the Resources You Use?, 93 4.1 Assessment, 101 Chapter 4 Assessment, 122 Earth and Human Activity: Science and Engineering Practices: Designing Solutions: Design to Reduce Waste, 729
Crosscutting Concepts	
(ES.ESS3.2.CCC.1) Scale Proportion and Quantity: Using concepts of orders of magnitude allows one to understand how a model at one scale relates to a model at another scale.	For supporting content, please see: SE/TE: Figure 4: Tar Sands Deposits, 97 Figure 5: Distribution of Oil Shale in the Green River Formation, 98 Nuclear Energy, 103-104 Map It! And Figure 21: Virgin Forests 1620-1992, 112
Performance Expectation	
(ES.ESS3.5) Construct a scientific explanation from evidence for how geological processes cause uneven distribution of natural resources.	SE/TE: Fossil Fuels, 95-96 Figure 2: U.S. Coal Fields, 95 Map It!, 95 Tar Sands and Oil Shale, 97 Figure 4: Tar Sands Deposits, 97 Figure 5: Distribution of Oil Shale in the Green River Formation, 98 Formation of Mineral Deposits, 98-100

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Disciplinary Core Ideas	
(ES.ESS3.5.DCI.1) Most elements exist in Earth's crust at concentrations too low to be extracted, but in some locations-where geological processes have concentrated them-extraction is economically viable.	SE/TE: Minerals, 45 How Minerals Form, 45-46 Fossil Fuels, 95-96 Tar Sands and Oil Shale, 97 Formation of Mineral Deposits, 98-100 Nonmetallic Mineral Resources, 100-101 Land Resources, 111-112 Layers Defined by Composition, 233-234 Discovering Earth's Composition, 237
Science and Engineering Practices	
(ES.ESS3.5.SEP.1) Constructing Explanations: Construct an explanation based on valid and reliable evidence obtained from a variety of sources (including students' own investigations, models, theories, simulations, peer review) and the assumption that theories and laws that describe the natural world operate today as they did in the past and will continue to do so in the future.	SE/TE: The Process of Science, 23 Scientific Knowledge, 24 4.2 Assessment, 107 4.4 Assessment, 116 Earth & Its Resources: Bingham Canyon, Utah: The Largest Open-Pit Mine, 117 & TE Notes Chapter 4 Assessment, 122 Uniformitarianism, 336
Crosscutting Concepts	
(ES.ESS3.5.CCC.1) Cause and Effect: Empirical evidence is required to differentiate between cause and correlation and make claims about specific causes and effects.	For supporting content, please see: SE/TE: TE, Build Reading Literacy, 111 4.3 Assessment, 112 4.4 Assessment, 116 Inquiry: Exploration Lab: Finding the Product that Best Conserves Resources, 118-119 Chapter 4 Assessment, 122

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